

Physics 141.
Lecture 22.



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Physics 141.
Lecture 22.

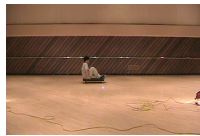
- Course Information.
- Entropy.
 - Quick review of topics discussed during our last lecture.
 - Heat capacity.
 - The Maxwell-Boltzmann velocity distribution of gas molecules.

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Analysis of experiment # 5.
Updated Timeline.

- ✓ 11/14: collisions in the May room
- ✓ 11/20: analysis files available.
 - <https://www.pas.rochester.edu/~tdimino/phy141/lab05/>
- ✓ 11/20: each student has determined his/her best estimate of the velocities before and after the collisions (analysis during regular lab periods).
- ✓ 11/21: complete discussion and comparison of results with colliding partners and submit final results (velocities and errors).
- ✓ 11/25: results will be compiled, linear momenta and kinetic energies will be determined, and results will be distributed.
- 12/4: office hours by lab TA/TIs to help with analysis and conclusions.
- 12/6: students submit lab report # 5.

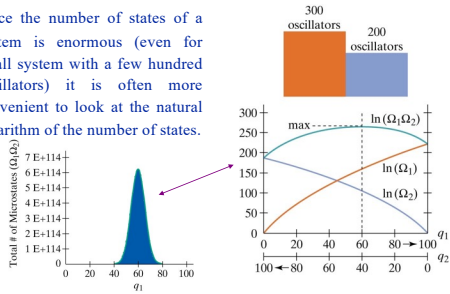


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Distributing energy. Number of states: different representations.

- Since the number of states of a system is enormous (even for small system with a few hundred oscillators) it is often more convenient to look at the natural logarithm of the number of states.



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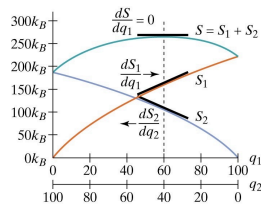
Entropy.

- Since the evolution of the system depends on the number of microstates, the concept of entropy is introduced.
- The entropy S of one of the blocks we have been discussing is defined as

$$S = k \ln \Omega$$

where k is the Boltzmann constant (1.4×10^{-24} J/K).

- The entropy of the total system is the sum of the entropy of the blocks that make up the system.



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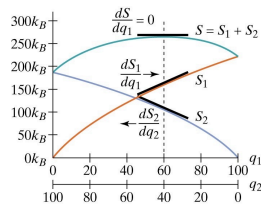
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Entropy and the Second Law of Thermodynamics.

- To achieve thermal equilibrium, the system will maximize its entropy.
- The most likely evolution of the system is the focus of the **second law of thermodynamics**:

If a closed system is not in equilibrium, the most probable consequence is that the entropy of the system will increase.

- Note: even when the two blocks are in thermal equilibrium, there may still be exchange of energy between the blocks, but the time averaged energy exchange will be zero.



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Entropy and temperature.

- The maximum entropy is reached when

$$\frac{dS}{dq_1} = \frac{dS_1}{dq_1} + \frac{dS_2}{dq_1} = 0$$
 or

$$\frac{dS_1}{dq_1} = -\frac{dS_2}{dq_1} = \frac{dS_2}{dq_2} = \frac{dS_2}{dq_1}$$
- Assuming that each oscillator has the same energy quantization, this condition is equivalent to

$$\frac{1}{T_1} = \frac{1}{T_2} \Leftrightarrow T_1 = T_2$$

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The Boltzmann distribution.

- Question: What does the fundamental theory of statistical mechanics tell us about the probability to see a single atom with a specific energy?
- Answer: A lot.
- Consider a single oscillator in contact with 300 other oscillators.
- The entropy of this system is equal to the entropy of the 300 oscillators.
- If the energy of the single oscillator increases, the number of states decreases.

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The Boltzmann distribution.

- The entropy curve is close to linear at small ΔE and its slope is equal to dS/dE_1 .
- We can approximate the dependence of the entropy on ΔE in the following manner:

$$S = A - \frac{dS}{dE} \Delta E = A - \frac{\Delta E}{T}$$

- The number of states of the system, as function of ΔE , can now be determined:

$$\Omega = e^{\ln \Omega} = e^{S/k} = e^{A/k} e^{-\Delta E/kT}$$

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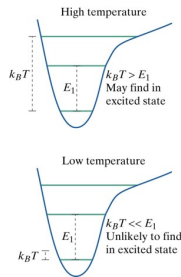
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The Boltzmann distribution.

- Since the probability is proportional to the number of states, we conclude that:

the probability of finding a microscopic system to be in a state with energy ΔE above the ground state of the system is proportional to $e^{-\Delta E/k_B T}$.

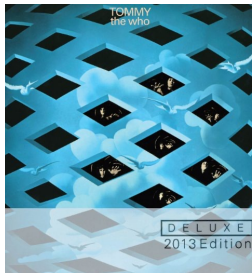
- This probability distribution is called the **Boltzmann distribution**.



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3 Minute 03 Second Intermission.



- Since paying attention for 1 hour and 15 minutes is hard when the topic is physics, let's take a 3 minute 03 second intermission.

- You can:
 - Stretch out.
 - Talk to your neighbors.
 - Ask me a quick question.
 - Enjoy the fantastic music.
 - Solve a WeBWork problem.



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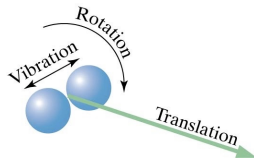
The Boltzmann distribution.

- The energy of a gas molecule can have various components:

- Translational
- Vibrational
- Rotational
- Gravitational

- For each component, the probability will show an exponential dependence on energy.

- I will first focus on the translational component.



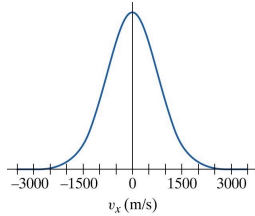
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The Boltzmann distribution.

- Consider that the gas molecule is moving with a velocity v_x along the x axis.
- The energy of the gas molecule, associated with its motion along the x axis, will be $(1/2)Mv_x^2$.
- The probability of finding the gas molecule with a velocity between v_x and $v_x + dv_x$ is equal to

$$P(v_x)dv_x \propto e^{-\left(\frac{1}{2}Mv_x^2\right)/kT} dv_x$$



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The Boltzmann distribution.

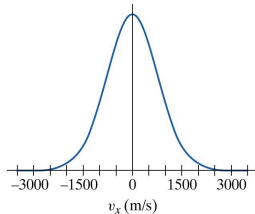
- We can do the same calculation for the velocity distribution along the y and z axes:

$$P(v_y)dv_y \propto e^{-\left(\frac{1}{2}Mv_y^2\right)/kT} dv_y$$

$$P(v_z)dv_z \propto e^{-\left(\frac{1}{2}Mv_z^2\right)/kT} dv_z$$

- These equations can be combined to determine the dependence of the probability on v:

$$P(v)dv \propto e^{-\left(\frac{1}{2}M(v_x^2+v_y^2+v_z^2)\right)/kT} dv_x dv_y dv_z = 4\pi v^2 e^{-\left(\frac{1}{2}Mv^2\right)/kT} dv$$



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The Maxwell-Boltzmann speed distribution.

- The probability to find a molecule with a speed between v and $v + dv$ is thus proportional to

$$P(v)dv \propto 4\pi v^2 e^{-\left(\frac{1}{2}Mv^2\right)/kT} dv$$

- But the probability to find the molecule with a speed between 0 and infinity is equal to 1. Thus:

$$1 = \int_0^{\infty} P(v)dv = 4\pi C \int_0^{\infty} v^2 e^{-\left(\frac{1}{2}Mv^2\right)/kT} dv = C \left(\frac{2\pi kT}{M} \right)^{3/2}$$

- The velocity distribution is thus equal to

$$P(v) = 4\pi \left(\frac{M}{2\pi kT} \right)^{3/2} v^2 e^{-\left(\frac{1}{2}Mv^2\right)/kT}$$

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The Boltzmann distribution and internal energy.

- The root-mean-square kinetic energy associated with the three translational degrees of freedom is thus equal to

$$\bar{K}_{rms} = \bar{K}_{x,rms} + \bar{K}_{y,rms} + \bar{K}_{z,rms} = \frac{3}{2}kT$$

- It turns out that the average energy associated with each degree of freedom, including vibrational and rotational degrees of freedom, is $(1/2)kT$.

- Note:**

- The internal energy only depends on the temperature; it does **not** depend on the mass of the gas molecules.
- At a given temperature, the rms velocity of heavier molecules will be smaller than the rms velocity of lighter molecules.

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The Boltzmann distribution and internal energy.

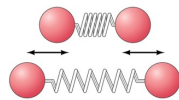
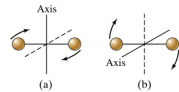
- Up to now we have assumed that the internal energy U of a gas is equal to $(3/2)kT$.

- This is correct for a monatomic gas, but is not correct for diatomic or triatomic gases.

- Predictions for a diatomic molecule:**

- Linear motion (3 degrees of freedom):
 $U = (3/2)kT$.
- Rotational motion (2 degrees of freedom):
 $U = kT$.
- Vibrational motion (2 degrees of freedom):
 $U = kT$.

- The number of degrees of freedom excited depend on the temperature.



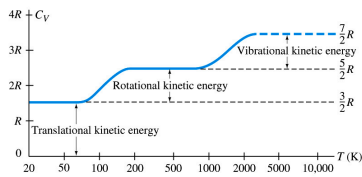
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The Boltzmann distribution and internal energy.

Remember: $Q = C \Delta T$

But: $K = (1/2)nkT$ and $\Delta K = (1/2)nk \Delta T$

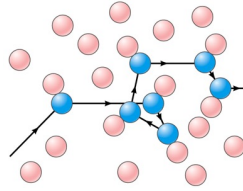


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Molecular speed and its consequences. The mean-free path.

- The RMS velocities of individual gas molecules are large. For example, for hydrogen at room temperature, the RMS velocity is 1920 m/s.
- Despite the large RMS velocity, the average diffusion velocity is much smaller and is largely determined by the mean-free path of the molecules.
- We expect that the mean-free path is inversely proportional to the cross-sectional area of the molecules and inversely proportional to the density.



Typical values of the mean-free path are between 10^{-8} and 10^{-7} m

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Molecular speed and its consequences. The propagation of sound.

- Typical values for v_{rms} :
 - For H at 300 K: $v_{rms} = 1,920$ m/s.
 - For N at 300 K: $v_{rms} = 517$ m/s.
- The speed of sound in these two gases is 1,350 m/s for H and 350 m/s for N.
- Note: The speed of sound in a gas will always be less than v_{rms} since the sound propagates through the gas by disturbing the motion of the molecules. The disturbance is passed on from molecule to molecule by means of collisions; a sound wave can therefore never travel faster than the average speed of the molecules. Since v_{rms} increases with T , we expect v_{sound} to increase with T .

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Molecular speed and its consequences. The composition of the atmosphere.

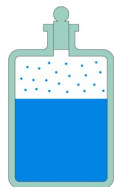
- Typical values for v_{rms} :
 - For H at 300 K: $v_{rms} = 1,920$ m/s.
 - For N at 300 K: $v_{rms} = 517$ m/s.
- A significant fraction of the H molecules in the atmosphere have a velocity that exceeds the escape velocity of the earth. This is one of the reasons that our atmosphere contains virtually no hydrogen and helium.
- Since the escape velocity decreases with decreasing planetary mass, the escape velocity of the moon is so low that it has no atmosphere (even the heavier gas molecules would quickly escape).

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Molecular speed and its consequences. Evaporation.

- Evaporation: transformation from the liquid to the gas phase.
- A microscopic view of evaporation:
 - Molecules with high velocity moving close to the surface can overcome the strong attractive forces between the molecules and escape from the liquid (evaporation).
 - The average velocity of the molecules left behind in the liquid will be lowered.
 - Since the average velocity is proportional to the temperature, the temperature of the liquid is lowered when evaporation takes place.



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Enough physics for this week!



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