Physics 141. Lecture 12.



1

Physics 141. Lecture 12.

- Course Information
- Chapter 8, Energy Quantization:
 Quantization of energy.
 Emission and absorption spectra.
 Vibrational energy levels.
 Rotational energy levels.
 Example of quantization.

 - Incoherent and coherent emission of light.

Frank L. H. Wolfs

Department of Physics and Astronomy, University of Rochester, Lecture 12, Page 2

2

Physics 141. Course Information.

- Homework:
 Homework set 5 is due on Friday 10/11 at 12 pm.
 No homework during the week of fall break.
 Homework set 6 is due on Friday 10/25 at 12 pm.

- Homework set 6 is due on Friday 10/25 at 12 pm.

 Laboratory:
 Laboratory # 3 took place on Monday 10/7.
 Laboratory # 3 office hours will be on Monday 10/21.
 Laboratory # 3 report will be due on Wednesday 10/23 at noon.

 Exam # 2:
 Thursday 10/17 at 8 am in Hoyt.
 I will review the material covered on Exam # 2 on Thursday 10/10.

- Next week:
 No recitations on Tuesday 10/15.
 Regular office hours and recitations on Wednesday 10/16 to help you prepare for Exam # 2.
 No office hours on Thursday 10/17

Chapter 8. Energy quantization.

- Until the early 20th century, scientists assumed that the total energy of a system could have any arbitrary value (constraint only by the potential energy of the system).
 - The energy of a planet in a circular orbit around the sun will have an energy that depends on the radius of the orbit. By changing the radius, we can change the energy of the system.
 - The success of the planetary model led to applications of this model to the description of atoms where electrons carry out orbital motion around the nucleus.
 - This model predicts that the total energy of the electron is only constraint by the electrostatic potential energy.

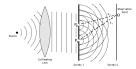
Frank L. H. Wolfs

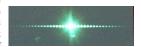
Department of Physics and Astronomy, University of Rochester, Lecture 12, Page

4

Quantization of light.

- Before 1905, light was assumed to be an electromagnetic wave, characterized by an amplitude and a wavelength.
- When the intensity of the light increases, the amplitude of the electromagnetic wave increases.
- Most phenomena could be understood in terms of the wave nature of light. Diffraction and interference provided compelling evidence that the wave picture of light was correct.





Frank L. H. Wolfs

Department of Physics and Astronomy, University of Rochester, Lecture 12, Page

5

Energy quantization. Quantization of light.

- But the photoelectric effect (emission of electrons from metals) could not be understood in terms of the wave model of light.
- To explain the photoelectric effect, light must be described in terms of "particle" properties. When the intensity of a "particle" beam increases, the number of particles increases but the energy of each particle does not change.
- The particles of light are called **photons**. The energy of a photon is determined by its wavelength (the typical energy for a photon of visible light is around 2 3 eV).
- The energy of the photon is hc/λ , where h is Planck's constant (= 6.6×10^{-34} Js) and λ is the wavelength of the light.

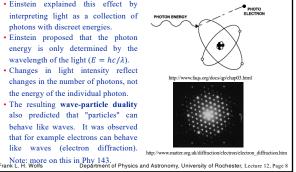
Frank L. H. Wolfs

Quantization of light. The photoelectric effect. · Experiments showed that light can liberate electrons from a material (the photoelectrons). following surprising observations were made: Photoelectric effect No electrons are liberated if the wavelength of the light is larger than some critical value (independent of the intensity of the light). The maximum energy of the electrons liberated does not depend on the intensity of the light. It only depends on the wavelength of the incident light.

7

Quantization of light. The photoelectric effect.

- · Einstein explained this effect by interpreting light as a collection of photons with discreet energies.
- Einstein proposed that the photon energy is only determined by the wavelength of the light $(E = hc/\lambda)$.
- · Changes in light intensity reflect changes in the number of photons, not the energy of the individual photon.
- The resulting wave-particle duality also predicted that "particles" can behave like waves. It was observed that for example electrons can behave like waves (electron diffraction).



8

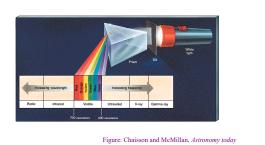
3 Minute 40 Second Intermission (42 seconds more than during lecture 11).



- Since paying attention for 1 hour and 15 minutes is hard when the topic is physics, let's take a 3 minute 40 second intermission.
- You can:
 - Stretch out.
- Talk to your neighbors.
- Ask me a quick question. Enjoy the fantastic music. Go asleep, as long as you wake up in 3 minutes and 40 seconds.

Frank L. H. Wolfs

Quantization of light. Different wavelengths, different energies.



10

Energy quantization. Energy levels of hydrogen.

Emission pattern of Hydrogen. "Astronomy! A Brief Edition," J. B. Kaler, Addison-Wesley, 1997.



- Light is emitted by heated hydrogen gas only at specific wavelengths.
- Excited hydrogen atoms can thus only emit specific amounts of energy.
- This implies that the energy states of hydrogen atoms are quantized.

Frank L. H. Wolfs

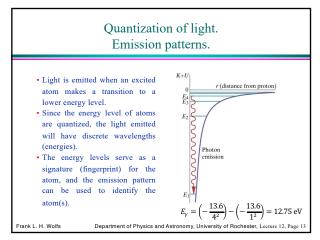
Department of Physics and Astronomy, University of Rochester, Lecture 12, Page 11

11

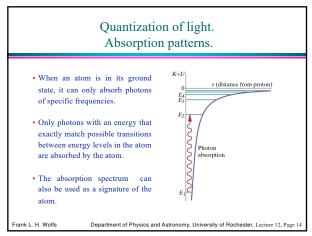
Energy quantization. Energy levels of hydrogen.

- Experiments examining the details of atomic structure showed that the electrons can only occupy certain specific energy levels.
- The measurements show that the total energy of the electron, K + U, is equal to $-13.6/N^2$ eV where N is an integer (N = 1, 2, ...).
- The quantization of the energy levels is a result of quantummechanical effects (to be discussed in more detail in Physics 143).

(+U	Unbound energies not quantized
0	r (distance from proton)
E ₄ E ₃	Bound-state energies quantized
	$E_N = K + U_{\text{elec}} = -\frac{13.6 \text{ eV}}{N^2}$
	$U_{\rm elec}$
E ₁	



13

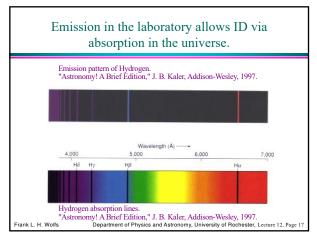


14

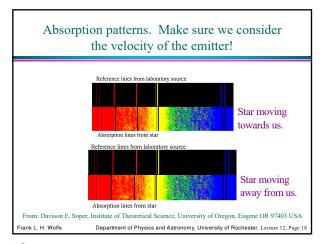
Quantization of light. Absorption patterns. • The atoms that absorb the photons will decay back to the E_2 Emitted E_6 light E_6 E_1 E_2 E_2 E_3 E_4 E_6 ground state. • When these transitions occur, the E_1 emitted photons will have the Detector same energies as the energies of E_5 E_6 E_6 E_6 E_6 E_6 E_5 the photons being absorbed. • However, the reemitted photons $E_7 \sim \sim \sim$ Incoming E₆ are emitted in random directions, and only a few are emitted in the direction of the incoming photons. Frank L. H. Wolfs Department of Physics and Astronomy, University of Rochester, Lecture 12, Page 15

Quantization of light. Absorption patterns. At room temperature, almost all hydrogen atoms will be in their ground state (E = -13.6 eV). • The first excited state in hydrogen is at E = -3.4 eV and a 10.2 eV photon is required to make the transition. · Since visible light has energies Photon absorption around 2 - 3 eV, the hydrogen will appear to be transparent to visible light. • Note: visible light can be absorbed when the atom is already in an excited state, but this is unlikely at room temperature. Frank L. H. Wolfs Department of Physics and Astronomy, University of Rochester, Lecture 12, Page 16

16

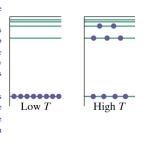


17



Absorption patterns. Temperature dependence.

- The absorption pattern depends on the population of states in the sample.
- · Atoms with excited states populated are able to absorb lower-energy light (longer wavelength) since the energy spacing between atomic levels decreases with increasing energy.
- Since the population patterns depends on temperature, we expect to see a temperature dependence of the absorption spectrum.



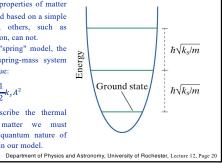
19

Other examples of quantization. The atomic model of matter.

- · Although many properties of matter can be understood based on a simple "spring" model, others, such as thermal conduction, can not.
- In the classical "spring" model, the energy of the spring-mass system can have any value:

$$E = \frac{1}{2} k_s A^2$$

• In order to describe the thermal properties of matter we must incorporate the quantum nature of these oscillators in our model.



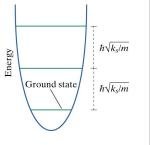
20

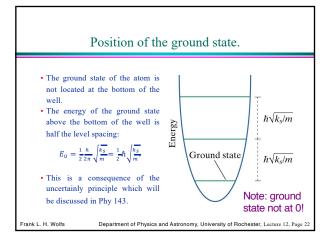
Other examples of quantization. The atomic model of matter.

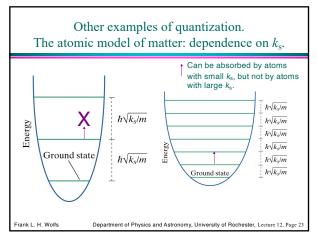
- The effect of the quantum treatment of the oscillator is that the energy of the system is quantized.
- It turns out that in many cases a potential well consistent with the spring force is consistent with the observed atomic properties.
- The energy levels in this well are quantized and the spacing between the

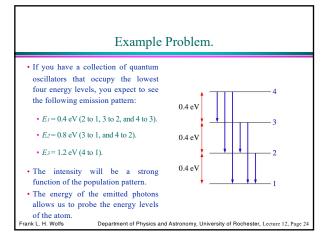
 $\Delta E = \frac{h}{2\pi} \sqrt{\frac{k_S}{m}} = \hbar \sqrt{\frac{k_S}{m}}$

• One important consequence of this quantization is that atoms can only transfer energy to each other in discrete amounts.









Other examples of quantization. Molecular vibrational energy. • When we measure the vibrational MWWWWW energy levels for a two-atomic molecule we find that at low energies the vibrational model works fine (nearly uniform energy spacing).

• At higher energies, the potential well starts to deviate from a harmonic oscillator well, and the vibrational energy levels are no longer uniformly spaced.

Nearly uniform energy spacing

25

Other examples of quantization. Molecular rotational energy.

- Molecules can also rotational energy.
- The rotational energy of a molecule is also quantized, but the spacing between levels increases with increasing excitation energy.
- The rotational energy is found to be equal to

 $E_\ell = \frac{1}{2I}\ell(\ell+1)\hbar^2$

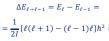
where ℓ is an integer $(0, 1, \dots)$ and I is the moment of inertia

(depends on mass and shape).
Frank L. H. Wdlfs Department of Physics and Astronomy, University of Rochester, Lecture 12, Page 24

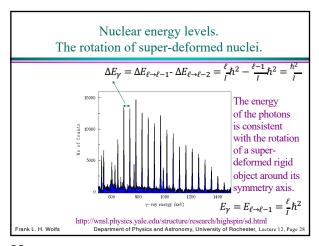
26

Other examples of quantization. Molecular rotational energy.

- When a molecule is in a rotational excited state, it will decay back to the ground state by gradually lowering its excitation
- The transition energies expected to be seen are



 $= \frac{1}{2I} \{ \ell(\ell+1) - (\ell-1)\ell \} \hbar^2 =$



Different systems. Different energies. State Energy level spacing (eV) Hadronic 100,000,000 Nuclear 1,000,000 Atomic 1 Molecular (vibrational) 0.01 Molecular (rotational) 0.0001 Frank L. H. Wolfs Department of Physics and Astronomy, University of Rochester, Lecture 12, Page 29

