Today in Astronomy 102: black hole evaporation

- □ Hawking's "area increase" theorem.
- Entropy and the area of the event horizon: the thermodynamics of black holes.
- Quantum-mechanical vacuum fluctuations and the emission of light by black holes: "Hawking radiation."
- □ Evaporation of black holes.
- □ Exotic matter.

Image: a typical intrauniverse wormhole. From <u>Matt Visser</u>, Victoria U. of Wellington



Results Exam # 2



Correlation between practice exam 2 and exam 2.



The horizon area theorem

In 1970, Stephen Hawking used general relativity to prove a useful rule called the **horizon area theorem**:

The total horizon area in a closed system containing black holes never decreases. It can only increase or stay the same.

Increases in total horizon area come from growth of black holes by collapse or accretion of "normal" matter, and by the coalescence of black holes.

Illustration (next page): a closed-off part of the universe. As time goes on, the total area of all the horizons in this closed system increases, owing to the growth of black holes by collapse, accretion and coalescence.

Illustration of Hawking's horizon-area theorem



So why should that seem strange?

- The horizon-area theorem is simple (and intuitively obvious), but represents a puzzle at a deeper level when one reflects on the **heat** and **disorder** in the matter that forms or falls into a black hole.
 - Before: the matter is hot, and there are lots of particles sharing the heat among themselves in the form of their random motions. A complete description of the system would thus have different entries for position and velocity for each particle – a vast number of *numbers* required.
 - □ After: the system can be completely described by only three numbers, its mass, spin and charge. It's orderly!
- The problem is, in all other natural processes **matter is never seen to go from a disorderly state to an orderly one all by itself.** This is in fact a law of thermodynamics...

Entropy and the second law of thermodynamics

Compare these two statements:

The **horizon area theorem**:

□ The total horizon area in a closed system never decreases. The **second law of thermodynamics**:

□ The total entropy of a closed system never decreases.

Entropy = the logarithm of the number of ways all of the atoms and molecules in a system can be rearranged without changing the system's overall appearance. A larger entropy means the system is more disorderly, or more "random."
Do black holes really have entropy as low as they seem to? Does horizon area have anything to do with entropy?

For adepts, scientists and, engineers

Others may ignore this page (i.e. **not on exams or HW**): Officially, the definition of entropy includes an additional multiplicative factor:

 $S = k \log \Omega$

where S = entropy,

- $k = \text{Boltzmann's constant}, 1.381 \times 10^{-16} \text{ erg K}^{-1},$
- Ω = number of equivalent rearrangements,

and by "log" is meant the natural (base *e*) logarithm.
For AST 102: suffice it to say that entropy is some number times the logarithm (any base) of the number of equivalent rearrangements.

Example of entropy: toys in a playroom (Thorne, pg. 424)

Extremely orderly: 20 toys on 1 tile



Number of equivalent rearrangements = 1; entropy = 0. This playroom floor has 100 tiles, on which the kids can arrange 20 different toys. Parents prefer the toys to be kept in an extremely orderly configuration, with all the toys piled on one tile in one corner, as shown. There is only one such arrangement; the entropy of this configuration is thus the logarithm of 1, which is zero.

Entropy in a playroom (continued)

Orderly: 20 toys on 10 tiles



Number of equivalent rearrangements = 10²⁰; "entropy" = 20. Parents might even accept this somewhat less orderly configuration: 20 different toys on 10 specific tiles. But there are lots of different equivalent arrangements (e.g. swapping the positions of two toys on different tiles produces a different arrangement that's still acceptable): 10^{20} of them, in fact, for an entropy value of $\log(10^{20}) = 20.$

Entropy in a playroom (continued)

Disorderly: 20 toys on 100 tiles



Number of equivalent rearrangements = 10⁴⁰; "entropy" = 40. Of course, when the kids are done playing, the floor looks like this: 20 different toys spread randomly over 100 tiles. There are $100^{20} = 10^{40}$ different ways to do that; the entropy is 40. And kids are like natural physical processes. Through their agency the room will not get less random; the entropy of the room full of toys and kids never decreases by itself.

Many consider the second law of thermodynamics to be the physical law least likely ever to be broken.

This is because its consequences are so easily tested experimentally that you can see it verified all around you every day. Among the important corollaries to the second law:

- Heat never flows by itself from a lower temperature to a higher temperature.
- Decreasing the temperature of one part of a closed system requires raising the temperature in other parts.
- The mechanical work (*organized* energy) that can be done by a heat engine is always less than the available heat (*disorganized* energy).
- □ Irreversible processes: two sequences energetically allowed, but only one sequence ever happens.

Lecture 19

Aside: irreversible processes and the Arrow Of Time

The laws of conservation of energy, momentum, and spin are indifferent to the order of time, and if they had the only say, all mechanical processes would be time-reversible.

□ For instance, a diver can bound from a springboard, fall 5 meters, and hit the water with kinetic energy about 4×10¹⁰ erg.



<u>Greg Louganis</u>, 3 m springboard, 1984 Summer Olympics

□ In the water he is brought to a halt, his kinetic energy converted to heat: this raises the temperature of the water (assumed uniform) by about 4.5×10⁻⁷ degrees C.

Irreversible processes and the Arrow Of Time

- So in principle it's possible for the pool to cool down by a tiny amount, transfer that energy to the diver in the form of kinetic energy, and spit him back out of the pool.
- □ As you know, that never happens: this process is **irreversible**.



<u>Greg Louganis</u>, 3 m springboard, 1984 Summer Olympics

The difference: the diver's kinetic energy is organized, the heat in the pool is disorganized. That divers never are flung back out of the pool is an expression of the fact that closed systems never become more organized spontaneously.

Irreversible processes and the Arrow Of Time

Thus a deep connection exists between the second law of thermodynamics and the direction of time, which brings us back to relativity:

From the standpoint of formal physics there is only one concept which is asymmetric in time, namely entropy. But this makes it reasonable to assume that the second law of thermodynamics can be used to ascertain the sense of time independently in any frame of reference; that is, we shall take the positive direction of time to be that of increasing entropy. (Panofsky and Phillips 1962).



Mid-lecture break (2 min. 56 sec.).

The Jets of NGC 1097 Image Credit & Copyright: Martin Pugh

Black holes form from large collections of atoms and molecules, with extremely large numbers of equivalent rearrangements (large entropy).

What happens to the entropy of this matter when it falls into a black hole? This was a burning issue in 1972.

□ Hawking (and the rest of the relativists in 1972): The entropy vanishes. Black holes and their horizons are extremely simple objects, with only one possible configuration ("equivalent rearrangement") each: that means zero entropy. Black holes therefore violate the second law of thermodynamics. That doesn't matter very much, because you can't get accreted material back out of the black hole anyway.

- Jacob Bekenstein (a Princeton graduate student in 1972, and all alone on this side of the argument): The second law of thermodynamics hasn't been violated in any other physical situation; why give up so soon? The entropy of the ingredients may be preserved, in a form proportional to the horizon area. If hole has entropy it also must have a temperature, which I find is proportional to the strength of gravity at the horizon.
 Hawking *et al.* (1972): But that would mean that the horizon is a black body at non-zero temperature that
 - horizon is a **black body** at non-zero temperature that obeys the laws of thermodynamics. Any such body must radiate light - as the hot filament in a lightbulb does, for instance. *Nothing can escape from a black hole horizon*; how can it radiate?

- □ **Hawking** *et al.* **(1972):** This contradiction implies that black holes cannot have entropy or temperature, and that they must violate the second law of thermodynamics.
- Bekenstein (1972): I can't think of any way for light, or anything else, to escape from a black hole; I admit that black holes can't radiate. But there must be something wrong with your viewpoint, because it must be possible for black holes to obey the laws of thermodynamics.

And, actually, Zel'dovich already had thought of a way for horizons to radiate light, a year previously but unknown to these contestants until several years later.

Hawking (1974): Oops.

- There is a way for black holes to emit radiation: it involves quantum-mechanical processes near the black hole's horizon.
 - The emission of light is exactly as one would expect from a black body with temperature that increases as the strength of gravity at the horizon increases.
 - Therefore the black hole has entropy, which increases as the area of the horizon increases.
 - Therefore black holes obey the laws of thermodynamics. **Bekenstein is right after all**.
- □ The way Hawking found includes the way Zel'dovich had previously found, as a particular case.

How black holes emit light and other particles

□ In subatomic matter, we see that energy conservation can be violated, though only temporarily and very briefly. This is expressed in one of Heisenberg's uncertainty principles:

$$\Delta E \times \Delta t \approx h \approx 10^{-27} \text{ erg-sec}$$

- □ Vacuum fluctuations: the shorter the time interval $\triangle t$, the larger the energy $\triangle E$ that can be **temporarily** produced. For extremely short time intervals, enough energy can be borrowed from the vacuum (i.e. nothingness) to produce photons, or even massive particles.
- □ The particles thus made are called **virtual particles**. They vanish again at the end of the time interval $\triangle t$.
- □ Virtual particles are produced as **particle-antiparticle pairs**.

How black holes emit light (continued)

Examples of particle-antiparticle pairs made from the vacuum:



How black holes emit light (continued)

Normally, virtual pairs vanish too quickly to be noticed, or to interact much with anything else.

Near a black hole horizon: what if one of the pair falls in, and the other doesn't, and is aimed just right? Gets away!



How black holes emit light (continued)

Details of black-hole emission, nowadays called **Hawking radiation**:

- □ Virtual pairs, produced by vacuum fluctuations, can be split up by the strong gravity near a horizon. Both of the particles can fall in, but it is possible for one to fall in with the other escaping.
- The escaping particle is seen by a distant observer as emission by the black hole horizon: black holes emit light (and other particles), though only in this weird way.
- □ The energy conservation "debt" involved in the vacuum fluctuation is paid by the black hole itself: the black hole's mass decreases by the energy of the escaping particle, divided by *c*². The emission of light (or any other particle) costs the black hole mass and energy.

Black hole evaporation

Hawking radiation is emitted more efficiently if the gravity at the horizon is stronger (i.e. its temperature is higher).

- Recall: horizon gravity is stronger for smaller-mass black holes.
- Thus an isolated black hole will eventually evaporate, as it radiates away all of its mass-energy. The smaller the black hole mass is, the larger the evaporation rate is.

The time it takes to evaporate:

- \Box 10⁹ M_{\odot} black hole: 10⁹⁴ years.
- \square 2 M_{\odot} black hole: 10⁶⁷ years.
- \Box 10⁸ gram black hole: 1 second (!)

Black hole evaporation (continued)

It's easy enough to derive the equation for the time it takes a black hole to evaporate *via* Hawking radiation that I usually assign it as a homework problem in Astronomy 142. The answer, for a black hole that starts off with mass *M*, is

$$t = \frac{10240\pi^2 G^2}{hc^4} M^3$$
$$= \left(\frac{8.407 \times 10^{-26} \frac{\sec}{gm^3}}{gm^3} \right) M^3 \quad ,$$

where $h = 6.626 \times 10^{-27}$ erg sec is Planck's constant, and you already know all the others.

(You need to understand, and learn how to use, this formula.)

Relativistic Heavy-Ion Collisions. Brookhaven National Laboratory, NY, USA



Relativistic Heavy-Ion Collisions. Brookhaven National Laboratory, NY, USA.





Creating Black Holes on Long Island.



Melville says he believes it would be microscopic at first but would grow exponentially, eventually obliterating Earth. "The black hole would first eat its way down toward the center of Earth and consume from the inside out. It would not be a good time to be around to see this. In the end ALL of Earth would be consumed."

When I started looking into this, I was stunned to find that other physicists are speculating along the same lines as Melville. The July 1999 Scientific American contains letters debating the possibility Melville raises, and the July 18 Sunday Times of London reported on and editorialized against the experiment, which it considers frighteningly dangerous. So it's not just paranoid physicists and rogue journalists concerned about the RHIC.

Hoping to forestall the end of the world, I contacted Brookhaven immediately. "We certainly do not wish to destroy the earth," sniffed spokeswoman Diane Greenberg, who clearly has been fielding plenty of questions like mine. Then she sent me a statement by Brookhaven Lab Director John Marburger, entitled "On Consequences of RHIC Operations."

"The amount of matter involved in the RHIC collisions is exceedingly small - only a single pair of nuclei is involved in each collision," Marburger states. "Our universe would have to be extremely unstable in order for such a small amount of energy to cause a large effect. On the contrary, the universe appears to be quite stable against releases of much larger amounts of energy that DARCHIV occur in astrophysical processes.

Read Fred "RHIC collisions will be within the spectrum of Moody's past energies encompassed by naturally occurring cosmic columns radiation. The earth and its companion objects in our solar system have survived billions of years of cosmic ray collisions with no evidence of the instabilities that have been the subject of speculation in connection with RHIC?

Playing at God

Why am I not reassured by this? The short answer is that the experiment is conducted by human beings - the same folks who brought you the internal combustion engine, which threatens to destabilize the planet's climate, and powerful antibiotics, which ultimately created an invincible staphylococcus bacterium. In other words, technopride goeth before the fall.

The longer answer is that Melville's scenario is perversely seductive in a Kubrickian sort of way. Think of Dr. Strangelove and 2001: A Space Odyssey. There are few things quite as persuasive as the vision of humans, their thirst for knowledge and progress insatiable, stumbling on a way to destroy the planet. It is an end-ofthe-world scenario that has launched a thousand movie

considers frighteningly dangerous.

WHAT DO YOU THINK? Should potentially dangerous experiments, like he one at Brookhaven, be allowed to proceed? O Yes O No Submit Vote

Not a scientific poll: for entertainment only

> "For once, maybe once in the history of the universe, we can avoid THE END. Have a nice day." Physicist David Melville

scripts.

Human progress has always had a nasty habit of producing unintended consequences - usually because the prideful progenitors of progress insist on poohpoohing any possibility of danger. Now, in recreating the beginning of the universe, we are essentially playing at being God - an unforgivable offense, punishable, as tragedians in the Bible and other literature have prophesied for centuries, by annihilation.

The Doomsday Machine

This Doomsday scenario dovetails creepily with the speculation put forth by the late Carl Sagan in his book Cosmos. Sagan believed that we could never find evidence of life anywhere else in the universe because the pattern of evolution has been the same everywhere: Life begins and evolves through millions of years to the moment when it destroys itself. The nature of consciousness is such that evolution itself is a doomsday machine.

Sagan considered nuclear war the likeliest cause of destruction, but the creation of an annihilating black hole is more plausible. Not only does it explain the apparent absence of life anywhere else in the universe, it also explains the absence of any ruins of past civilizations. A black hole removes all traces of everything - including of the creating civilization's planet.

"So why am I telling you this?" Melville's message to me ends. "I think this should be brought out into the general public's view. For once, maybe once in the history of the universe, we can avoid THE END. Have a nice day."

Fred Moody is the author of I Sing the Body Electronic: A Year with Microsoft on the Multimedia Frontier and of The Visionary Position: The Inside Story of the Digital Dreamers Who Made Virtual Reality a Reality. His column appears on alternate Wednesdays.

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Ready When You Are	we can avoid THE END. Have a nice
	day." Physicist David Melville
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Lecture 19

Astronomy 102

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WEBLINKS Ion Collider Microcosmic Bang RHIC RHIC Experiments Physical Review Focus

Creating Black Holes on Long Island. Do not worry!

Brookhaven National Laboratory

ENERGY

Added Brookhaven Director

workings."

Marburger, "Nature has been creating

those at RHIC for billions of years, and

disaster related to those collisions. RHIC does r

brings a rare phenomenon into the view of our

instruments so we can puzzle out its inner

take us beyond the limits of natural phenomena. It

collisions of energies comparable to

there is no evidence of any kind of

person going up in an elevator feels slightly heavier, just as they would if gravity on Earth were stronger.

In the same way, the rapid deceleration of RHIC ions as they smash into each other for a very short period of time (about 10⁻²³ second) is similar to the extreme gravitational environment in the vicinity of a black hole. This means that RHIC collisions should emit a bunch of thermal particles similar to the "Hawking radiation" emitted by a black hole. Since Hawking radiation is the cause of black hole decay, not formation, its existence would be yet another reason that RHIC cannot produce a real gravitational black hole.

Eurther discussion by Kharzeev on RHIC and black holes...

Operated by Brookhaven National Laboratory, the Relativistic Heavy Ion Collider is sponsored by the U.S. Department of Energy Office of Science, Office of Nuclear Physics,

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Black Holes at RHIC?

Before RHIC began operations in 2000, some were concerned that it would produce black holes that would threaten the earth. Here's why those concerns were unfounded.

Technology

Committee Review of Speculative "Disaster Scenarios" at RHIC

Home RHIC Science News Images Videos For Scientists

In July 1999, Brookhaven Lab Director John Marburger convened a committee of distinguished physicists to write a comprehensive report on the arguments that address the safety of speculative disaster scenarios at RHIC. The scenarios are:

Creation of a black hole that would "eat" ordinary matter. Initiation of a transition to a new, more stable universe Formation of a "strangelet" that would convert ordinary matter to a new form.



search



>> View the committee report (PDF).

Physics of RHIC Discoveries STAR PHENIX PHOBOS & BRAHMS Electron-Ion Collider Accelerator Complex Why? The Future

Why RHIC Cannot Produce a Real Gravitational Black Hole

By Physicist Dmitri Kharzeev

Horatiu Nastase, a member of the high-energy physics theory group at Brown University, wrote a paper, posted on the preprint website arxiv.org, in which he claimed that collisions at the Relativistic Heavy Ion Collider could produce the analog of a black hole.

Horatiu is referring to a mathematical similarity between the physics of the real world, which govern RHIC collisions, and the physics that scientists use to describe a theoretical, "imaginary" black hole in a hypothetical world with a different number of space-time dimensions (more than the four dimensions --three space directions and time - that exist in our world). That is, the two situations require similar mathematical wrangling to analyze. This imaginary, mathematical black hole that Horatiu compares to the RHIC fireball is completely different from a black hole in the real universe; in particular, it cannot grow by gobbling up matter. In other words, and because the amount of matter created at RHIC is so tiny, RHIC does not, and cannot possibly, produce a true, star-swallowing black hole.

This does not mean, however, that RHIC cannot study some of the phenomena that happen in the vicinity of black holes, as explained in a paper we wrote with Kirill Tuchin, also of Brookhaven's theoretical nuclear physics group. The explanation for this begins with Einstein's "Equivalence Principle," which states that gravity and acceleration (or deceleration) are actually equivalent forces. The principle explains why a



does not, and cannot possibly, produce a true, star-swallowing black hole.



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http://www.bnl.gov/rhic/blackHoles.asp

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Astronomy 102

How long is that?

The Large Hadron Collider (LHC) in Switzerland **might**, **conceivably**, be able to produce tiny black holes, with mass 2.50×10⁻²⁰ gm. How long will they last before evaporating?

A. 1.31×10⁹ sec B. 1.31 sec C. 1.31×10⁻²⁴ sec D. 1.31×10⁻⁵⁴ sec E. 1.31×10⁻⁸⁴ sec

Hawking radiation is really tiny and usually insignificant compared to other sources of light commonly found near black holes.

That's why it takes massive black holes so long to evaporate. Consider how it works out for a 2 M_{\odot} black hole.

Luminosity, if accreting at its maximum (Eddington) rate:

$$L = 2.5 \times 10^{38} \text{ erg/sec} = 6.6 \times 10^4 L_{\odot}$$

Luminosity due to Hawking radiation:

$$L = 2.3 \times 10^{-22} \text{ erg/sec} = 5.9 \times 10^{-56} L_{\odot}$$

Exotic matter

Part of this story should strike you as weird: Why is it that the black hole can *consume* a particle, and wind up *decreasing* in mass and energy?

- Because in the strongly warped spacetime near the horizon, virtual particles made from vacuum fluctuations turn out to have negative energy density.
 - **Energy density** = energy per unit volume.
- These particles indeed have positive mass -- look at the one that escaped! -- but their mass is distributed very strangely over spacetime. (Quantum-mechanically speaking, particles have nonzero volume; this is an aspect of the wave-particle duality.)
- Matter with negative energy density is generally called exotic matter.

Exotic matter (continued)

Theoretical details of exotic matter, according to the present partial marriage of general relativity and quantum mechanics (incompletely known; only studied intensively since 1985):

- Quantum mechanical vacuum fluctuations in flat spacetime - far from any strong gravitational field always have zero net energy density; they can never be exotic.
- However, in warped spacetime, vacuum fluctuations are in general exotic: their net energy density is negative, according to a distant observer measuring the energy density by observation of the deflection of light by the ensemble of fluctuations. The stronger the curvature, the more negative the energy density looks.

Exotic matter (continued)

Results of **theoretical calculations** of deflection of light by a black hole, with and without vacuum fluctuations:



Done! In the Shadow of Saturn.

